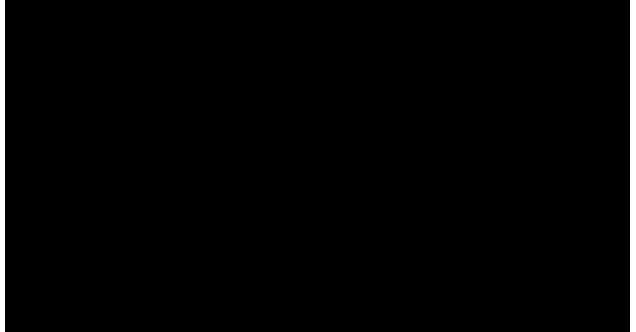
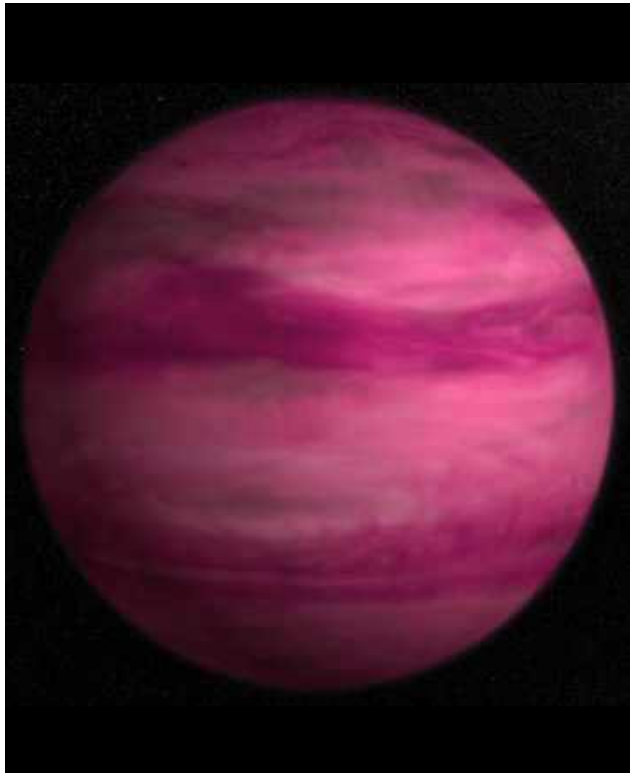
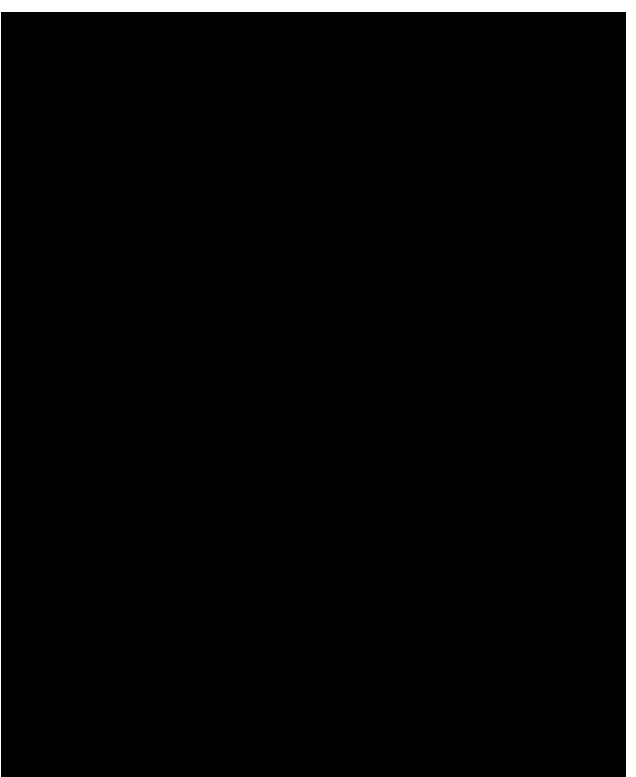
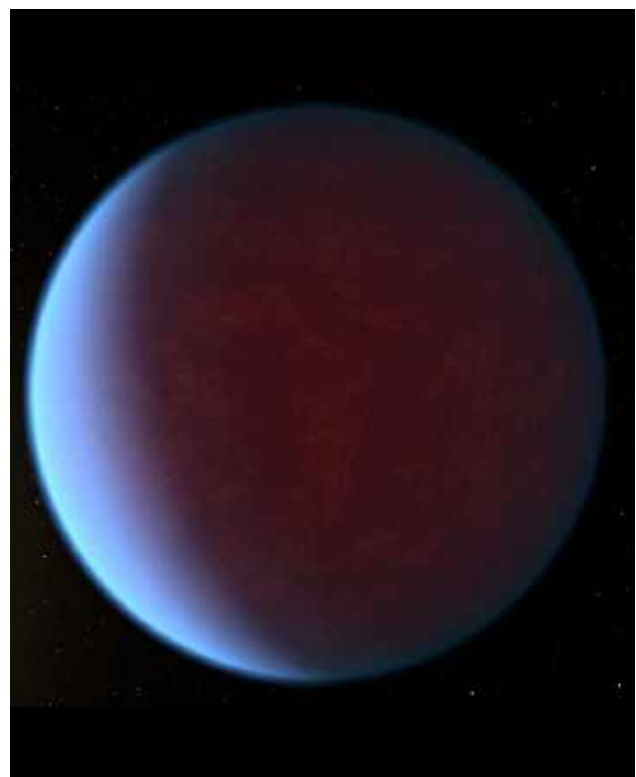


# PLANETARY YEARBOOK PHOTOS

STUDYING PLANET  
FORMATION IN SNAPSHOTS

Mackenna Wood

UNC Chapel Hill



# TABLE OF CONTENTS

- Background: What we already know
- Method: How to learn more
- New planets and what they tell us
- Questions?



# THE BUILDING BLOCKS OF A PLANET

1

DUST

2

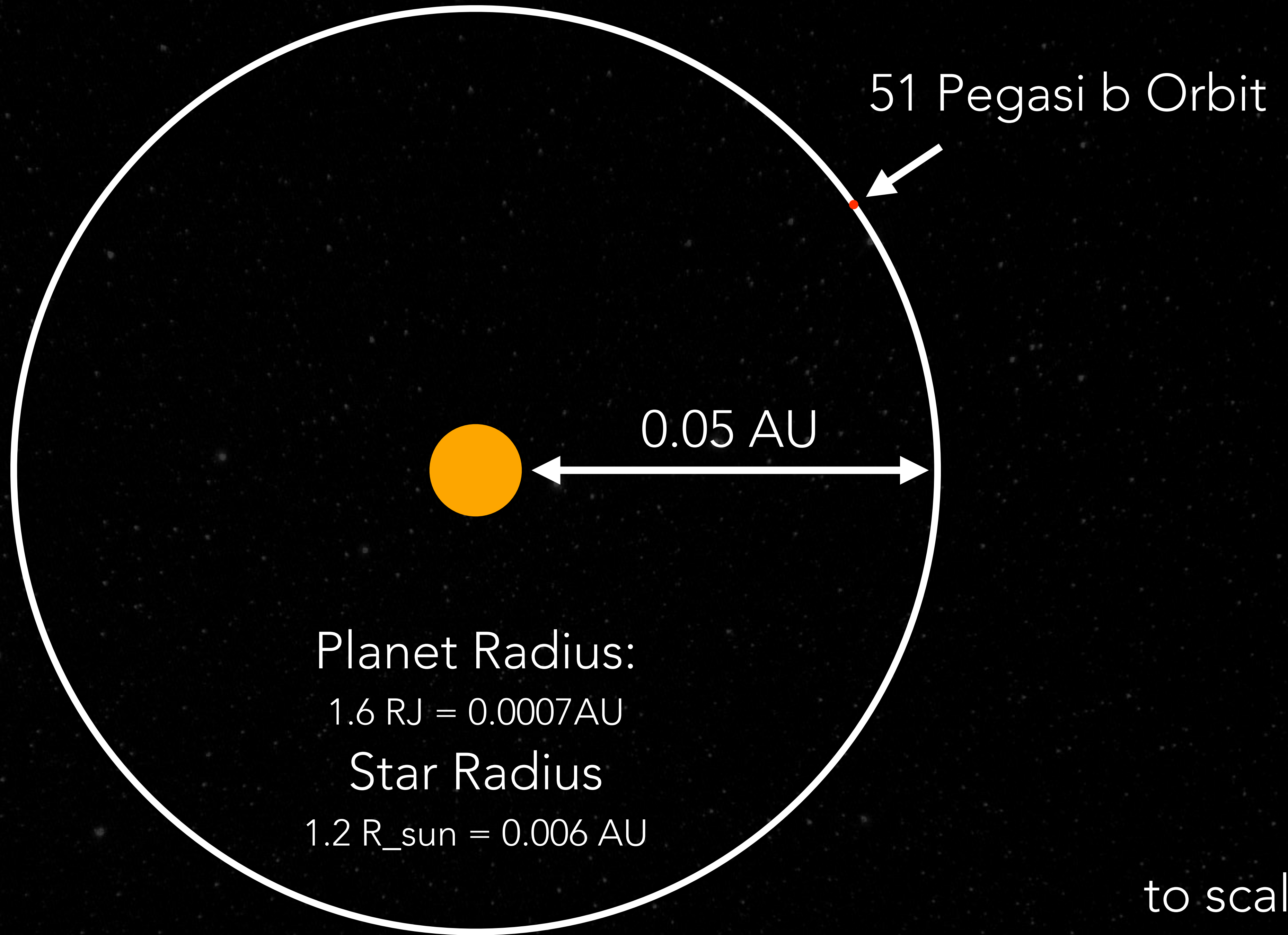
GRAVITY

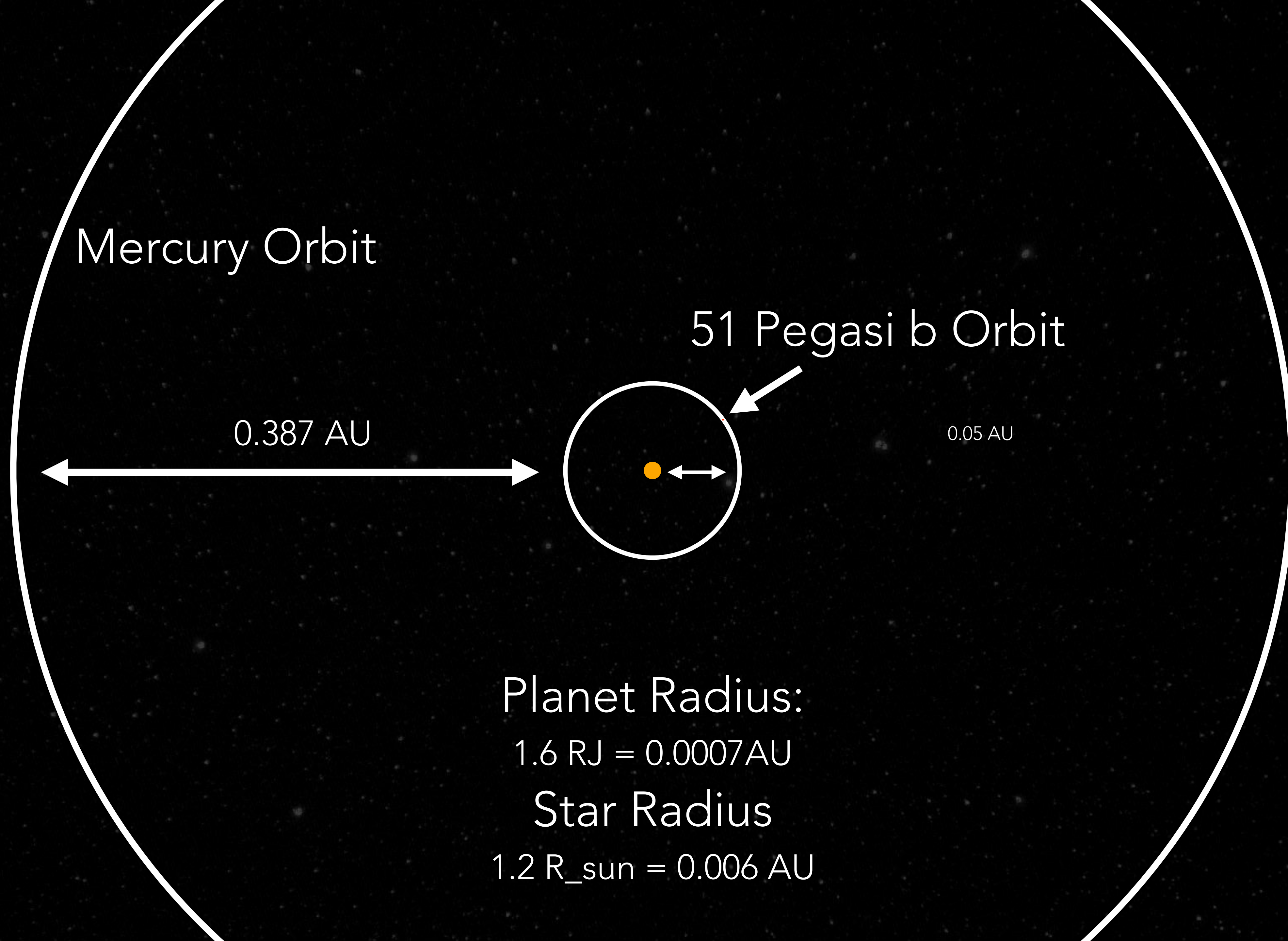
- A star forms from a cloud of dust
- The leftover material from the cloud stays in a disk around the star
- Small particles in the disk collide and stick together
- They stick to more particles, until the clump is big enough to attract even more
- The planetesimal picks up all the dust in its orbit
- The forming planets pick up all the dust in the disk



# WHAT WE ALREADY KNOW

- It's hard for planets to form extremely close to the star, because there isn't a lot of dust there
- But we find planets in that area, very close to their stars





to scale



51 Pegasi b  
Radius: 0.00076 AU



Mercury  
Radius: 0.000016 AU

to scale



# PLANET MIGRATION

- Method 1: Friction with the disk
- Method 2: Interactions with other planets

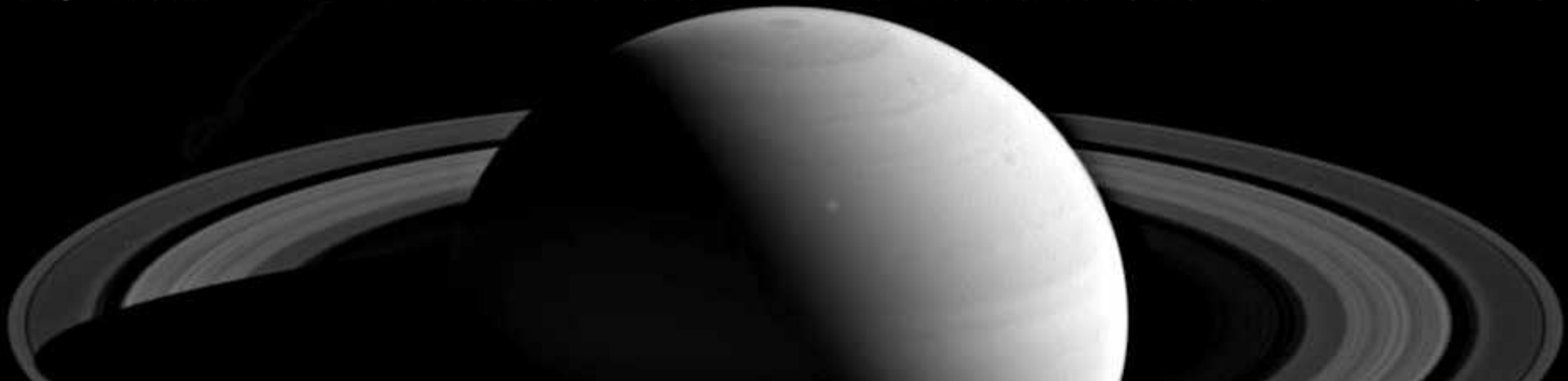
# CHANGING ATMOSPHERES

- Earth has gone through three atmospheres, only the last of which is capable of supporting human life
- Mars had an atmosphere and lost it
- Is it normal for planets to cycle through more than one atmosphere? When do they form and how?



# OTHER PLANET FEATURES

- Planets can have rings and moons
- Do these form with the planet?
- Do they last the whole life or do they come and go?



# REMAINING QUESTIONS

- How do planets change between their formation in the disk and their development into full 'mature' planets?
- Do all planets migrate? What causes them to migrate?
- Do all planets form atmospheres but lose them? Is it common for planets to have multiple atmosphere compositions during their lives?
- How do other planetary features change?

# THE GOAL

- We want to learn about how planets form and change over time.
- But the lifetime of a planet is much longer than the lifetime of all of modern astronomy, much less one astronomer
- So rather than watch one planet, we have to look at many planets and determine how they age from that



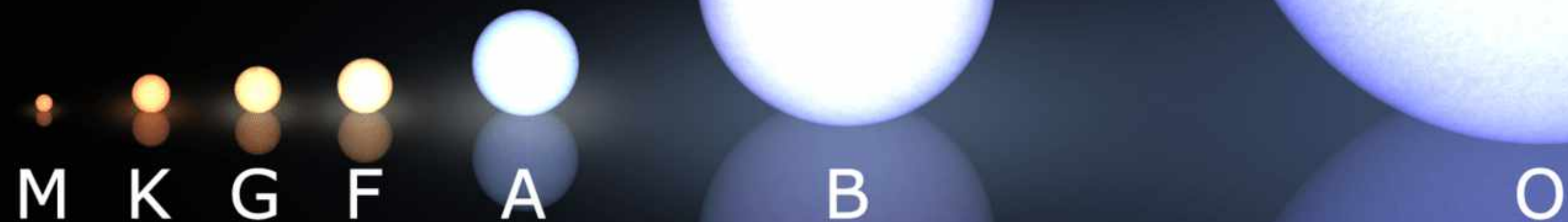


# SO HOW DO WE TELL THE AGES OF PLANETS?

- Planets have roughly the same ages as the stars they orbit
- Measuring the age of a star still has most of the same problems as measuring the age of a planet



- They evolve differently depending on their size
- They form as singles or as binaries or triples
- They form in different environments, giving them different compositions



# IT'S BETTER IN A GROUP

- Star clusters are groups of stars that all have the same age
- Because they have many stars age measurements of them are more accurate
- Then you know the age of any star or planet in the group



# WHAT IS A YOUNG PLANET?

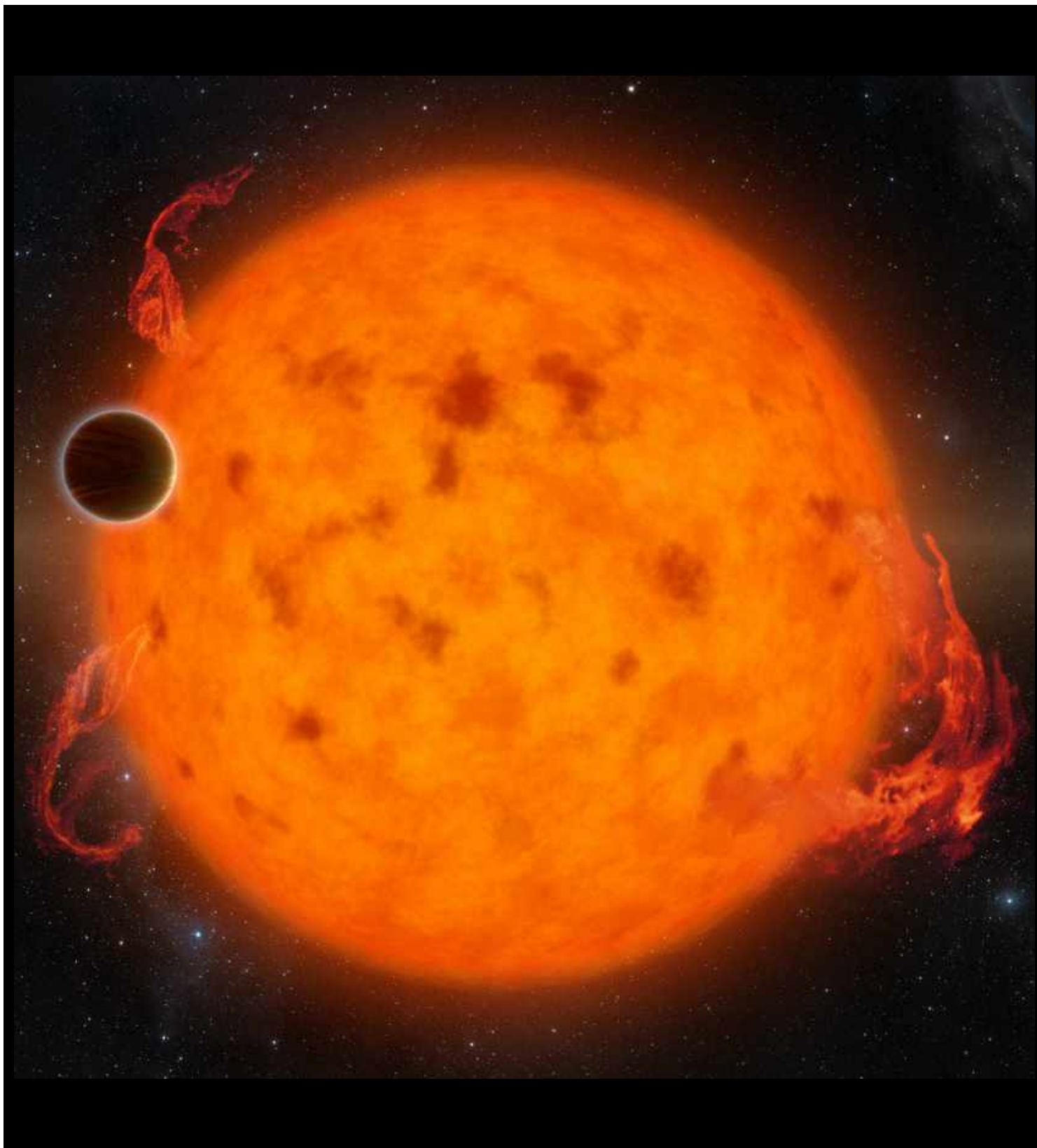
< 13 JUPITER  
MASSES

< 100 MILLION  
YEARS

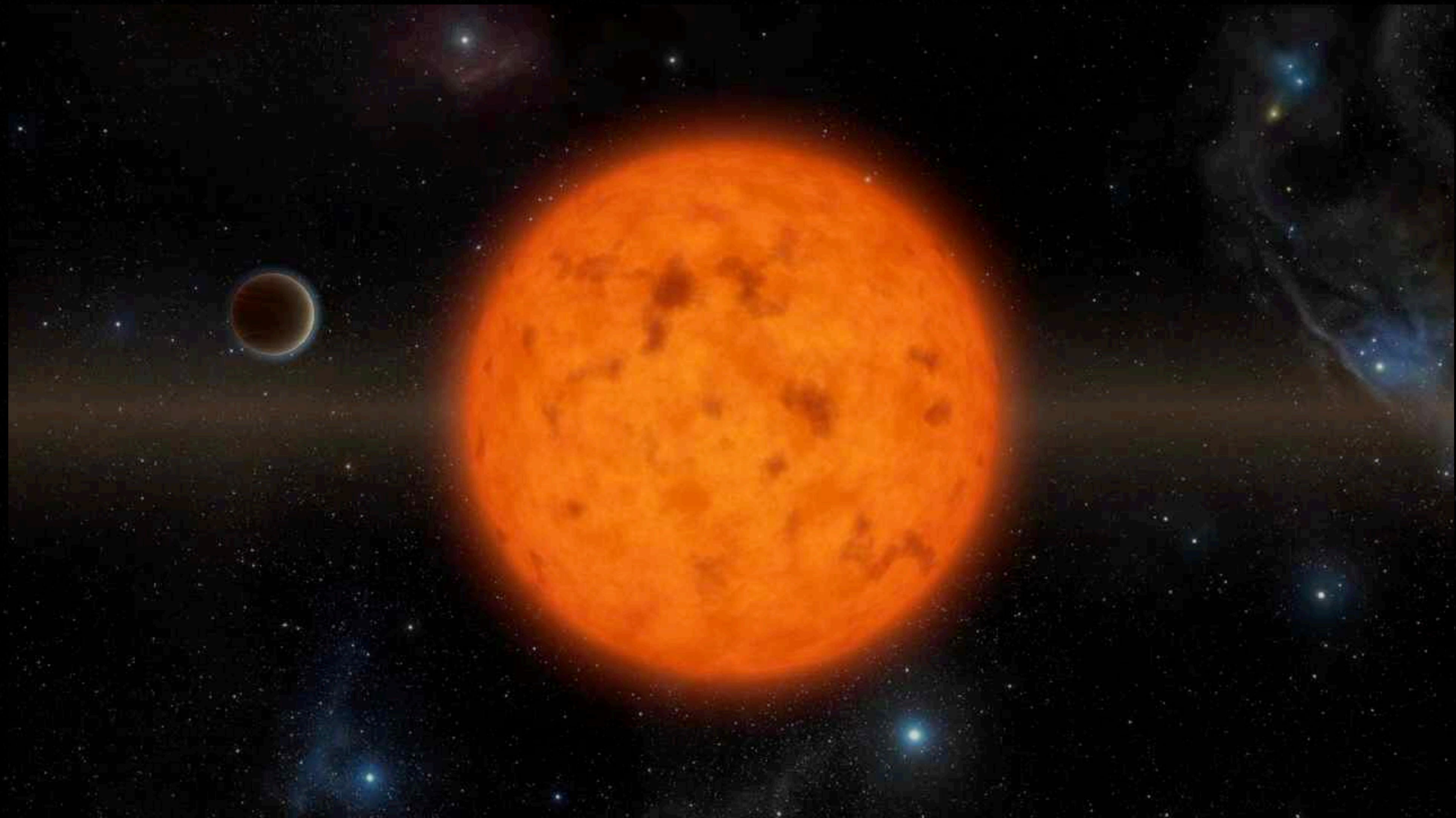
# KNOWN YOUNG PLANETS

- K2-33b ~ 10 Million Yrs
- V830 Tau b ~ 2 Million Yrs
- DS Tuc Ab ~ 45 Million Yrs
- V1298 Tau b ~ 23 Million Yrs

K2-33B



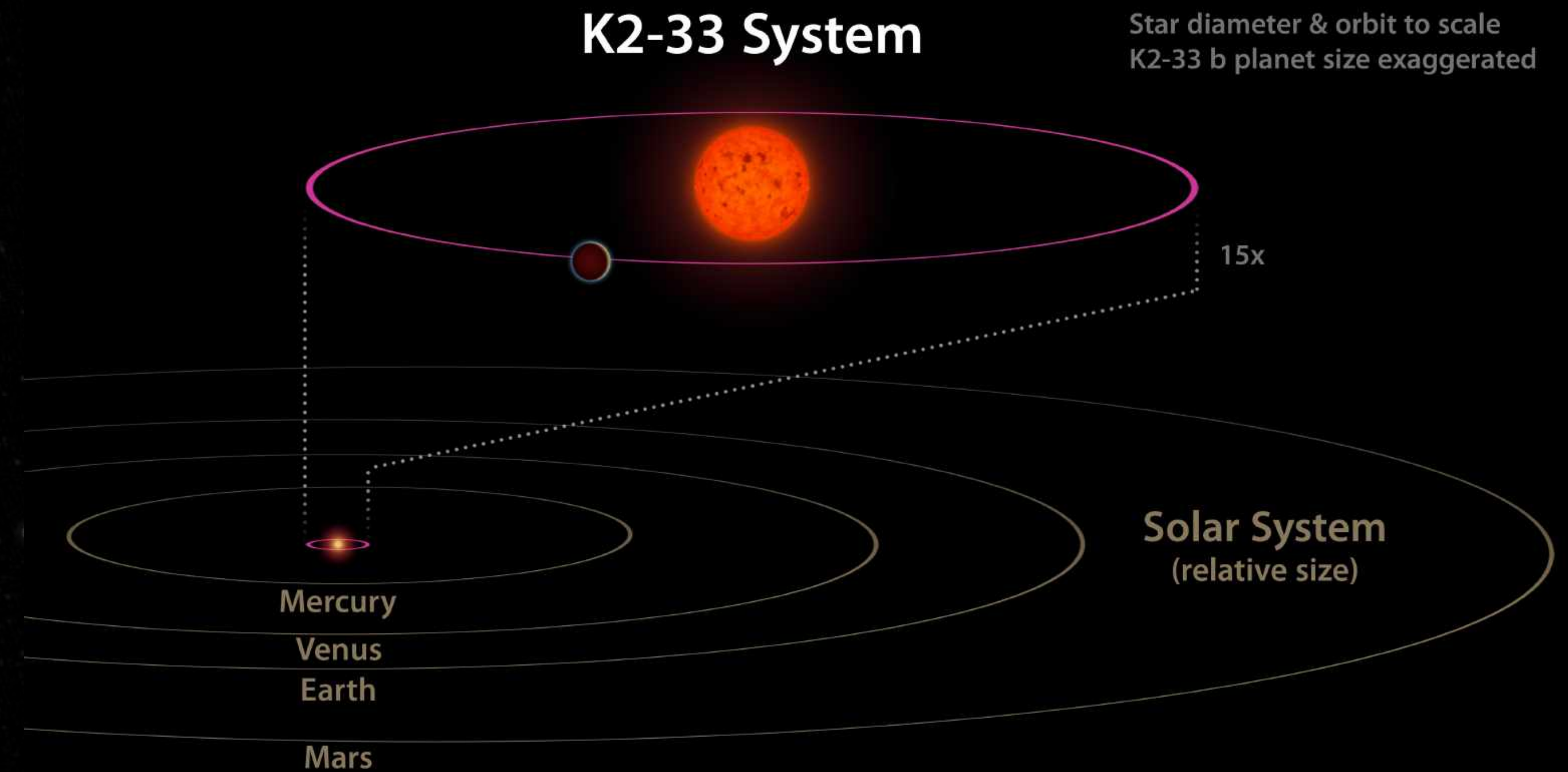
K2-33B



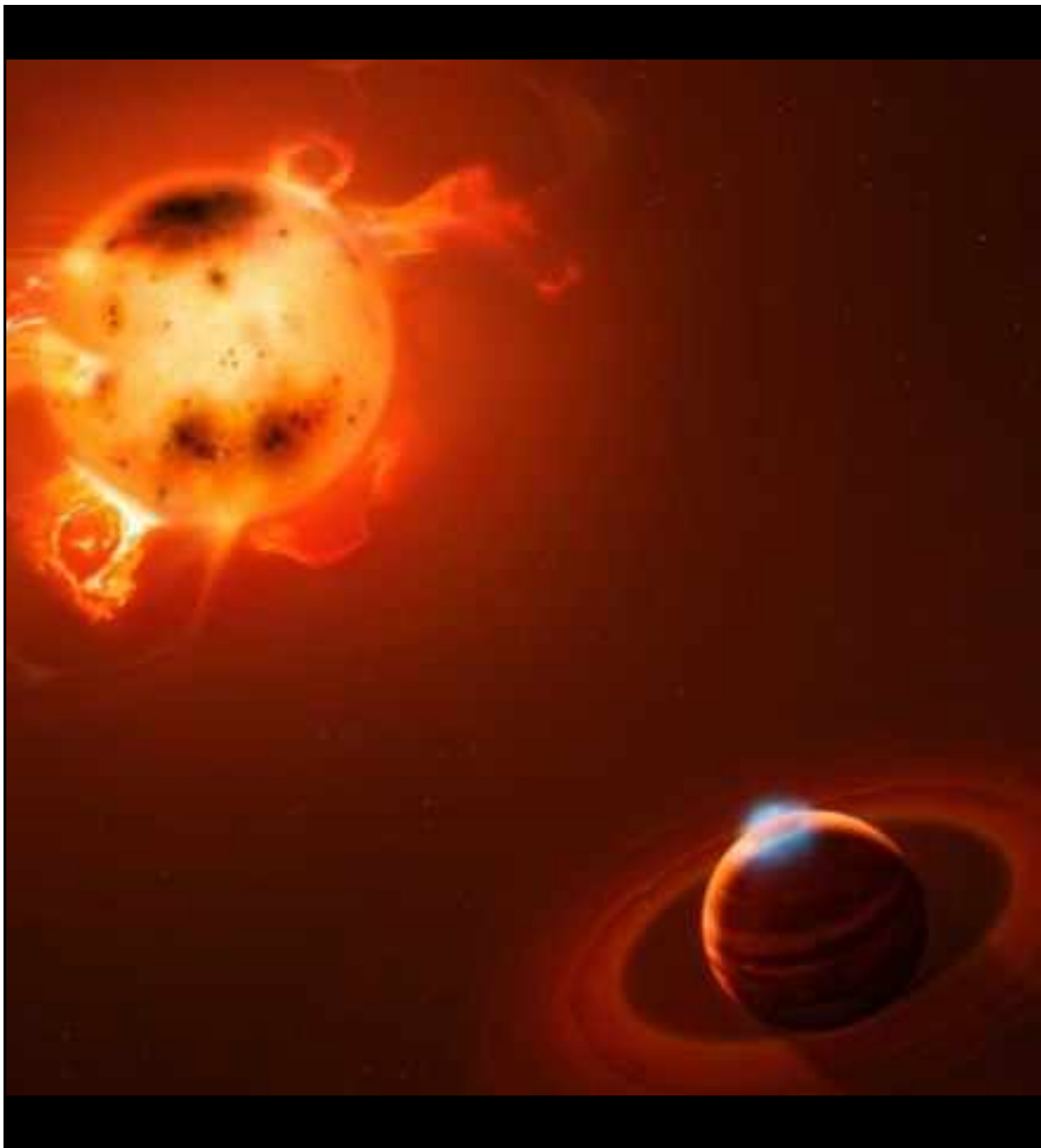
Movie Credit: NASA/JPL-Caltech

# K2-33B

- K2-33b is very close to its star
- It's only 11 million years old
- This means that the planet must have migrated inwards in the first 11 million years



# V830 TAU B





# V-830 TAU B

- V 830 Tau b is also very close to its star
- It's only 2 million years old

# DS Tuc AB AND V 1298 Tau B

- Provide useful laboratories for studying young planets

# CONCLUSION

- There are many things we don't know about how planets form and change
- To understand this we have to find many young planets of various types and find how each one fits into the puzzle
-

Questions ?